

# Estimation of sinking velocities using free-falling dynamically scaled models: foraminifera as a test case

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## 12 [Summary Statement](#)

13 We developed a novel method to determine the sinking velocity of biologically important micro-  
14 scale particles using 3D printed scale models.

## Abstract

The velocity of settling particles is an important determinant of distribution in extinct and extant species with passive dispersal mechanisms, such as plants, corals, and phytoplankton. Here we adapt dynamic scaling, borrowed from engineering, to determine settling velocities. Dynamic scaling leverages physical models with relevant dimensionless numbers matched to achieve similar dynamics to the original object. Previous studies have used flumes, wind tunnels, or towed models to examine fluid flows around objects with known velocities. Our novel application uses free-falling models to determine the unknown sinking velocities of planktonic foraminifera – organisms important to our understanding of the Earth’s current and historic climate. Using enlarged 3D printed models of microscopic foraminifera tests, sunk in viscous mineral oil to match their Reynolds numbers and drag coefficients, we predict sinking velocities of real tests in seawater. This method can be applied to study other settling particles such as plankton, spores, or seeds.

## Introduction

The transport of organisms and biologically derived particles through fluid environments strongly influences their spatiotemporal distributions and ecology. In up to a third of terrestrial plants (Willson et al., 1990), reproduction is achieved through passive movement of propagules (e.g., seeds) on the wind. In aquatic environments, propagules of many sessile groups from corals (Jones et al., 2015) to bivalves (Booth, 1983) are dispersed by ambient currents, eventually settling out of the water column to their final locations. Furthermore, most dead aquatic organisms (from diatoms to whales) sink, transporting nutrients to deeper water and contributing to long term storage of carbon (De La Rocha & Passow, 2007). In the case of microfossils, sinking dynamics of the original organisms even influences our reconstructions of the Earth’s paleoclimate (Van Sebille et al., 2015). Crucially, the horizontal distances over which all these biological entities are transported, and therefore their distributions, are affected by their settling velocities (Ali et al., 2011).

Measuring the individual settling velocities of small particles directly is challenging, especially when they are too small to be imaged easily without magnification (e.g. Walsby and Holland, 2006). Here we apply dynamic scaling, an approach commonly used in engineering, to circumvent this difficulty and accurately quantify the kinematics of sub-millimeter scale free-falling particles using enlarged physical models. We use scaled-up physical models in a high-viscosity fluid, enabling easy measurements of settling speed, orientation, and other parameters using inexpensive standard high-definition web cameras. While dynamically scaled models have previously been employed to study a number of problems in biological fluid mechanics (e.g. Vogel, Ellington and Kilgore, 1973; Vogel, 1987; Vogel, 1994; Koehl, 2003), the study of freely-falling particles of complex shape – for which settling speed is the key unknown parameter – presents a unique challenge to experimental design that we overcome in this work.

Engineering problems such as aircraft and submarine design often are approached using scaled-down models in wind tunnels or flumes to examine fluid flows around the model and the resulting fluid dynamic forces it is subjected to. To ensure that the behaviour of the model system is an accurate

representation of real life, similarity of relevant physical phenomena must be maintained between the two. If certain dimensionless numbers (i.e., ratios of physical quantities such that all dimensional units cancel) that describe the system are equal between the life-size original and the scaled-down model, “similitude” is achieved and all parameters of interest (e.g., velocities and forces) will be proportional between prototype and model (Zohuri, 2015). Intuitively, the model and real object must be geometrically similar (i.e., have the same shape), so that the dimensionless ratio of any length between model and original,  $Length_{model}/Length_{real}$ , is constant – this is the scale factor ( $S$ ) of the model. Less obvious is the additional requirement of dynamic similarity, signifying that the ratios of all relevant forces are constant. For completely immersed objects sinking steadily at terminal velocity (achieved quickly for most small particles, see Time to Terminal Velocity), dynamic similarity is achieved by matching the Reynolds number ( $Re$ ).

$Re$  is a measure of the ratio of inertial to viscous forces in the flow (Batchelor, 2000; within a biological context Vogel, 1994), and is typically defined as:

$$Re = \frac{L U \rho_{fluid}}{\mu}$$

where  $\rho_{fluid}$  is density of the fluid ( $\text{kg m}^{-3}$ );  $L$  is a characteristic length (m) of the object;  $U$  is the object’s velocity ( $\text{m s}^{-1}$ ); and  $\mu$  is the dynamic viscosity ( $\text{N s m}^{-2}$ , or  $\text{Pa s}$ ) of the fluid. In cases where  $L U \rho_{fluid}$  is large compared to  $\mu$ , e.g. whales, birds, and fish ( $Re \approx 3 \times 10^9 - 3 \times 10^6$ , Vogel, 1994), inertial forces dominate. In cases where  $L U \rho_{fluid}$  is relatively small compared to  $\mu$ , e.g. sperm, bacteria ( $Re \approx 3 \times 10^{-2} - 1 \times 10^{-5}$ ), viscous forces dominate. Finally, when  $L U \rho_{fluid}$  is of comparable magnitude to  $\mu$ ,  $Re$  is intermediate and one cannot discount either inertial or viscous forces. If the scaled model and original system exhibit identical  $Re$ , the relative importance of inertial versus viscous forces is matched between the two and any qualitative features of the flows (e.g. streamlines) will also be identical.

Dynamically scaled physical models exhibiting the same  $Re$  as the original systems have been used in a number of biological studies. Vogel and La Barbera (1978) outline the principles of dynamic scaling: to obtain the same  $Re$  when enlarging small organisms, the fluid flow must be slower and/or the fluid more viscous, and when making smaller models of large organisms, the fluid flow must be faster and/or the fluid less viscous. For instance, Vogel (1987) used air in place of water flowing at lower speeds when investigating the refilling of the squids mantle during swimming by scaling a model up 1.5 times relative to the animal’s actual size. More recently, Stadler et al (2016) investigated sand inhalation in skinks with 3D-printed enlarged models, using helium instead of air (thereby increasing viscosity) as the experimental fluid. Koehl and colleagues have studied crustacean antennule flicking (lobsters (Reidenbach et al., 2008), mantis shrimp (Stacey et al., 2002) and crabs (Waldrop et al., 2015)) as well as the movements of copepod appendages (Koehl, 1995) with enlarged models, using mineral oil in place of water. Finally, perhaps the largest change in scale was employed by Kim et al (2003), who modelled the bundling of *E. coli* flagella at a scale factor of  $\sim 61,000$ , submerged in silicone

oil ( $10^5$  times more viscous than water), and rotated at 0.002 rpm compared to the 600 rpm observed in real bacteria (Sowa & Berry, 2008).

In all the above studies, basic kinematics such as speeds in the original system were relatively easy to measure, and the experiments aimed to reveal the forces involved (e.g. hydrodynamic drag) or details of the fluid flow such as the pattern of streamlines. Since the representative speed  $U$  of the original system was known, designing experiments to achieve similitude was relatively straightforward because the  $Re$  was also known *a priori* – in these cases, the model size, speed, and working fluid properties were simply interrelated through  $Re$  (Eqn 1). For instance, once a working fluid and the model size were chosen, the required towing speed was obvious. However, in the case of sedimentation of small particles (e.g. spores, seeds, plankton), the sinking speed ( $U$ ) is the key unknown. With an unknown sinking speed, the operating  $Re$  is also unknown, so it is not straightforward to design experiments that achieve similitude with the original system. Here we present an iterative methodology leveraging 3D printed dynamically scaled models that allows determination of the sinking speeds of small objects of arbitrarily complex shape.

We use planktonic foraminifera as an example of a small (200 – 1500  $\mu\text{m}$ ) biological particle for which the settling velocity is important and typically unknown. Foraminifera are a phylum of marine ameboid protists (B. K. Sen Gupta (ed.) , 2002; Schiebel & Hemleben, 2005). By secreting calcium carbonate, foraminifera produce a multi-chambered shell (test) which, in planktonic foraminifera, can grow up to 1500  $\mu\text{m}$  in diameter, and which frequently exhibits complex shape (Table 1). Once the organism dies or undergoes reproduction, the empty test sinks to the ocean floor, and so oceanic sediments contain substantial numbers of foraminifera tests. Foraminifera account for 23-56% of the oceans production of carbonate ( $\text{CO}_3$ ) (Schiebel, 2002), an important factor in climate change models (Passow & Carlson, 2012). Of particular interest for climate predictions is calculating the flux of tests reaching the ocean floor (Schiebel, 2002; Jonkers & Kučera, 2015). While there are more than 30 extant species and over 600 species in the fossil record, settling velocities are known for only 14 species of foraminifera (Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014,  $3.41 \times 10^{-4}$  -  $6.8 \times 10^{-2} \text{ m s}^{-1}$ ,  $Re \approx 18 - 55$ ).

## Materials and Methods

### Similitude and Settling Theory

We assume that the size (i.e.,  $L$  – defined as the maximum length parallel to the settling direction,  $A$  – defined as the projected frontal area, and  $V$  – the particle volume not including any fluid-filled cavities), 3D shape ( $\Psi$ , here treated as a categorical variable due to our consideration of arbitrarily complex morphologies, see Table 1), and density ( $\rho_{\text{particle}}$ ) of the original sinking particle are known, while the sinking speed ( $U$ ) is unknown. The properties of the fluid surrounding the original particle (i.e.  $\rho_{\text{fluid}}$ ,  $\mu$ ) are also known, and our goal is to design experiments in which we sink a scaled-up model particle in a working fluid of known  $\rho_{\text{fluid}}$  and  $\mu$  in order to determine the model particle's sedimentation speed and, via similitude,  $U$  of the original particle.

While previous work (Berger et al., 1972; Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014) suggests that the  $Re$  of sinking forams should be  $10^0 - 10^2$ , the exact value of  $Re$  for morphology  $\Psi$  is assumed to be unknown. Hence, it is not immediately clear what size the model should be (i.e., the scale factor  $S = L_{model}/L_{real}$ ) in order to match this  $Re$  in the experiments and ensure similitude. Solving for both  $Re$  and  $S$  simultaneously requires additional mathematical relationships beyond Eqn 1.

Throughout, we use a superscript O to refer to the *Original* values of dimensioned variables at life size (e.g.,  $L^O$ ,  $V^O$ ,  $A^O$ ,  $\rho_{particle}^O$ ,  $U^O$ ) and  $Re^O$ ,  $C_D^O$  for the values of the dimensionless Reynolds number and drag coefficient (defined below) corresponding to real particles sinking in the original fluid (e.g., seawater of  $\rho_{fluid}^O, \mu^O$ ). While the fluid dynamics of flow around a particle of particular shape  $\Psi$  can be considered theoretically over a range of  $Re$ , only the dynamics at  $Re^O$  and  $C_D^O$  will represent the “operating point” corresponding to the life size particle settling speed  $U^O$ .

When a particle is sinking steadily at its terminal velocity, the sum of the external forces acting on the particle is zero (Eqn 2); that is, the upward drag force ( $F_{drag}$ , Eqn 3) and buoyant force ( $F_{buoyancy}$ , Eqn 4) must balance the weight of the particle ( $F_{weight}$ , Eqn 5):

$$\Sigma F = F_{drag} + F_{buoyancy} - F_{weight} = 0$$

where

$$F_{drag} = \frac{1}{2} C_D(\Psi, Re) \rho_{fluid} U^2 A$$

Eqn 3 introduces the drag coefficient  $C_D(\Psi, Re)$ , a dimensionless descriptor of how streamlined an object is. Both  $C_D$  and  $Re$  must be matched to achieve similitude.  $C_D$  depends on the shape of the object  $\Psi$ , including its orientation relative to the freestream flow – for instance,  $C_D$  of a flat plate oriented parallel to laminar flow is as low as 0.003 while  $C_D$  of a flat plate oriented perpendicular to the flow is  $\sim 2.0$  (Munson et al., 1994). However, in addition to object geometry,  $C_D$  also depends on qualitative characteristics of the flow, such as whether it is laminar or turbulent – that is,  $C_D$  also depends on  $Re$ .  $C_D$  of a sphere decreases from about 200 at  $Re = 0.1$  to about 0.5 at  $Re = 1000$ ;  $C_D$  generally decreases with  $Re$  for most shapes (Munson et al., 1994; Morrison, 2013). While  $C_D$  does not depend on object size directly, larger objects generally experience higher drag forces and this is captured by the inclusion of particle area ( $A$ ) in the expression for  $F_{drag}$  (Eqn 3). For brevity, we will omit  $\Psi$  hereafter and write the drag coefficient as  $C_D(Re)$ .

The buoyant force ( $F_{buoyancy}$ , Eqn 4) and weight ( $F_{weight}$ , Eqn 5) are both expressed using particle volume ( $V$ ), gravitational acceleration ( $g$ ), and density of the fluid ( $\rho_{fluid}$ ) or particle ( $\rho_{particle}$ ), respectively:

$$F_{buoyancy} = V \rho_{fluid} g$$

$$F_{weight} = V \rho_{particle} g$$

Substituting Eqns 3, 4, and 5 into 2 and eliminating  $U$  via the definition of  $Re$  (Eqn 1) yields an expression for the drag coefficient obtained through a force balance (indicated by a superscript  $\mathcal{F}$ ):

$$C_D^{\mathcal{F}}(Re) = \frac{2 (\rho_{particle} - \rho_{fluid}) V g}{\rho_{fluid} U^2 A} = \left( \frac{2 (\rho_{particle} - \rho_{fluid}) V g}{\rho_{fluid} A} \right) \left( \frac{\rho_{fluid} L}{Re \mu} \right)^2 = \frac{2 (\rho_{particle} - \rho_{fluid}) \rho_{fluid} V g L^2}{A Re^2 \mu^2}$$

Note that this expression can be simplified further upon identification of the dimensionless Archimedes number  $Ar = g L_{Ar}^3 \rho_{fluid} (\rho_{particle} - \rho_{fluid}) / \mu^2$  if the cubed length scale  $L_{Ar}^3 = VL^2/A$ , yielding  $C_D^{\mathcal{F}}(Re) = Ar/Re^2$ , as previously highlighted by others (e.g. Karamanev, 1996). However, we will proceed with the original form of Eqn 6 to keep key variables such as  $L$  explicit.

If  $C_D$  were known for a particular morphology, we could simply substitute values corresponding to the original test in seawater into Eqn 6 and solve for  $Re = Re^O$  and thus  $U^O$  via Eqn 1, immediately solving the problem of unknown settling speed. Unfortunately, the complex shapes of foraminifera (Table 1) coupled with the implicit dependence of  $C_D$  on  $Re$  means that both variables are generally unknown, and thus far we have only one constraining relationship between  $C_D$  and  $Re$ . More information is required to determine where along this constraint curve the operating  $C_D^O$  and  $Re^O$  are located. This information can come from experiments in which the sinking speeds of scaled-up model particles of various sizes (i.e., scale factors  $S$ ) in a viscous fluid are measured directly, allowing us to calculate  $Re$  via Eqn 1 and then  $C_D^{\mathcal{F}}(Re)$  via Eqn 6 for the models, with appropriate values substituted for each experiment. For clarity, we can rewrite Eqn 6 for a model in terms of  $S$  and the original test parameters ( $L^O, A^O, V^O$ ):

$$C_D^{\mathcal{F}}(Re) = \frac{2 (\rho_{particle} - \rho_{fluid}) \rho_{fluid} (V) g (S L^O)^2}{(S^2 A^O) Re^2 \mu^2}$$

where we use the fact that for a model,  $L = S L^O$  and  $A = S^2 A^O$ . While one would also expect  $V = S^3 V^O$  for 3D printed models, limitations of our 3D printer led to variation in  $V$  that we overcame using a more general empirical relationship between  $S$  and  $V$  based on mass measurements – see 3D Printer Limitations. Eqn 7 represents a constraining relationship between  $C_D$  and  $Re$  for the sinking particle, which we use to collect  $(Re, C_D)$  experimental data points at several  $S$ . Once sufficient data are collected, we can construct a new, empirical relationship (e.g., a cubic spline fit) between  $C_D$  and  $Re$  for a particular particle shape, which we term  $C_D^E(Re)$ . Finally, we can solve for the operating  $Re^O$ ,  $C_D^O$ , and  $U^O$  by finding the intersection point between the  $C_D^{\mathcal{F}}(Re)$  constraint curve specific to life-size particles sinking in seawater (i.e., Eqn 7 with  $S = 1$  and  $\rho_{particle}^O, \rho_{fluid}^O, \mu^O$ ) and our empirical  $C_D^E(Re)$

spline curve valid for a particular particle shape moving steadily through any fluid. MATLAB code can be downloaded from <https://github.com/matthewwalkerbio/Dynamic-scaling>.

## Study Species

To construct an empirical  $C_D^E(Re)$  curve for a particular test morphology, we started with 3D scans of individual specimens from 30 different species (Table 1). The majority of the species were selected from the University of Tohoku museum's database, eforam Stock (<http://webdb2.museum.tohoku.ac.jp/e-foram/>), with a micro computed tomography ( $\mu$ CT) scan resolution between 2.5 and 3.6 pixels per  $\mu$ m, and were exported as 3D triangular mesh (STL format) files. Specimens of an additional three species were scanned using synchrotron radiation based micro-computed tomography (SR $\mu$ CT). Imaging was performed at the Imaging Beamline P05 (IBL) (Greving et al., 2014; Haibel et al., 2010; Wilde et al., 2016) operated by the Helmholtz-Zentrum-Geesthacht at the storage ring PETRA III (Deutsches Elektronen Synchrotron – DESY, Hamburg, Germany). Specimens were imaged at an photon energy of 14 keV and with a sample to detector distance of 17 mm. For each tomographic scan, 900 projections at equal intervals between 0 and  $\pi$  were recorded. Tomographic reconstruction was done via a classical filtered back projection using the RECLBL library (Huesman et al., 1977). For processing, raw projections were binned two times resulting in an effective pixel size of the reconstructed volume of 1.44  $\mu$ m. These scans were segmented and rendered using SPIERS (Sutton et al., 2012), and again exported in STL format and are available from [Morpho Source](#). Meshes of all foraminifera were manually checked in Meshlab (Callieri et al., 2012) for integrity.

For species where more than one scan was available, the scan that contained the best-preserved specimen was chosen. By only including one specimen per species, this approach neglects phenotypic plasticity which is demonstrated in planktonic foraminifera (e.g. Lohmann, 1983; Morard *et al.*, 2013), but was chosen due to limitations of  $\mu$ CT scan availability and time constraints on the project.

## 3D printing and model preparation

The 3D scans allowed us to easily fabricate scaled-up (scale factor  $S$ ) physical models of each specimen using a FormLabs Form1+ (Formlabs, Somerville, Massachusetts, USA) 3D printer, using FormLabs Clear Resin Version 2 (Formlabs, Somerville, Massachusetts, USA) with a layer thickness of 50  $\mu$ m (see **Error! Reference source not found.**D for examples) and x-y resolution of 200  $\mu$ m. Models were washed and flushed with isopropanol to remove excess resin following Formlabs' guide and allowed to air dry. Support material was removed (**Error! Reference source not found.**D.vi), and the models lightly sanded with 400 grit Wet 'n' Dry paper, followed by a final isopropanol wash to remove any remaining residue. Once dry, models were filled with mineral oil in preparation for sinking. Clear resin was chosen to allow each model to be checked for bubbles (which would increase the buoyancy of the model). Any bubbles were removed using a 30-gauge needle and syringe.



Following convention when defining the area  $A_{particle}$  used in the definition of  $C_D$  (Eqn 3), we measured projected area of the sinking foraminifera. Referring to high resolution images of the sinking model (**Error! Reference source not found.D**), a digital model of the foraminifera was manually aligned to measure the projected area in a plane perpendicular to the sinking direction (**Error! Reference source not found.D**). We used the same procedure to measure the maximum length parallel to the flow ( $L$ ) for the calculation of  $Re$  (**Error! Reference source not found.D**). These choices facilitated objective comparisons of  $C_D$  across morphologically diverse species, to be detailed in a future study.

### 3D Printer Limitations

Whilst in principle, the volume of a printed model should simply scale according to  $V = S^3 V^0$ , due to inherent limitations of the 3D printer as well as difficulty in removing excess resin from small models, we found that this expectation was usually not satisfied, and weighing the models showed that  $M/\rho_{particle} > S^3 V^0$  where  $M$  is particle mass (**Error! Reference source not found.C**). Therefore, we estimated  $V$  of each model by weighing on an Entris 224-1S mass balance ( $\pm 0.001$  g) and assuming  $\rho_{particle}$  was  $1121.43 \pm 13.73$  kg m<sup>-3</sup>, based on the average mass of five 1 cm<sup>3</sup> cubes of printed resin. Furthermore, whenever a predicted value for  $V$  at a given scale factor  $S$  was needed, i.e. in Eqn 7 (see Remaining iterations), we based this on cubic spline interpolation of our  $V(S)$  data for existing models when sufficient data were available, with extrapolation based on cubic scaling of  $V(S)$  if required (see **Error! Reference source not found.C**):

$$V^{predict}(S) = \begin{cases} H(S), N \geq 3 \\ S^3 V^0, N < 3 \text{ or } S = 1 \end{cases}$$

where  $N$  is the number of existing volume measurements and  $H$  represents the cubic spline fit of  $V$  vs  $S$ . Note that because we always directly measured  $V$  by weighing after printing each model, and it is not necessary to achieve the exact  $Re$  and  $C_D$  of the operating point ( $Re^0$  and  $C_D^0$ ) in the experiments (see Remaining iterations), the empirical spline-based volume prediction was not strictly required for our method to succeed. It merely aids in improving the rate of convergence of our iterative approach by reducing the difference between our anticipated and actual  $Re$ ,  $C_D$  for each experiment.

### Settling tank

The models were released in a cylindrical acrylic tank (0.9 m in diameter and 1.2 m in height) of mineral oil ("Carnation" white mineral oil, Tennants Distribution Limited, Cheetham, Manchester, UK;  $\rho = 830$  kg m<sup>-3</sup>,  $\mu = 0.022$  Pa S) filled to a depth of 1.18 m (approximately 750 L). The tank was fitted with a custom net and net retrieval system (**Error! Reference source not found.A**) to allow easy retrieval of the models after their descent, allowing each model to be sunk 5 times. Integrated to the net retrieval system was the release mechanism, which was held centrally over the tank, with the grasping parts submerged below the oil level. This ensured that each model was released in a controlled and repeatable fashion.

## Particle imaging

To minimise reflections, the tank was surrounded by a black fabric tent-like structure. This also served as a dark background to facilitate visualisation of the model during descent. The tank was illuminated with a single 800 lumen LED spotlight placed underneath the tank and, as the Formlabs' Clear Resin is UV-fluorescent, two 20W "Blacklight" UV fluorescent tubes were placed above the tank.

The sinking models were recorded using two Logitech C920 HD webcams (Logitech, Lausanne, Switzerland), placed at 90° to each other (**Error! Reference source not found.A**) and recording at 960 pixels x 720 pixels and ~30 frames per second, allowing monitoring of the position and orientation of the particle in 3D as it fell. As these consumer-grade webcams use a variable frame-rate system, a custom MATLAB script was used to initiate camera recording, recording both frames and frame timestamps. Videos were recorded for 500 frames (~17 s). Sinking velocity was calculated over a central 0.8 m depth range, ensuring the model was at terminal velocity (see Time to Terminal Velocity) whilst also avoiding end effects which could slow the model as it reached the bottom of the tank. Based on observations of suspended dust, there was no discernable convection in the tank during any trials that might potentially affect sinking velocities. The curved walls of the tank introduced distortion, which was removed using the MATLAB toolbox "Camera Calibrator" (Mcandrew, 2004). Pixel size was 1.06 pixels per millimetre with a mean reprojection error of 0.5 pixels, therefore distance measurements (for calculating sinking velocities) were accurate to within 0.5 mm (0.06% of the traversed depth).

## Velocity calculation

Models were tracked in distortion-corrected frames using a modified version of Trackbac (Guadayol et al., 2017; Guadayol, 2016). The per-frame centroid coordinates obtained were then paired with the timestamp values recorded to calculate average settling velocity components in 2D for each camera (below,  $U_x$  is horizontal speed from camera one,  $U_y$  is horizontal speed from camera two, and  $U_{z,1}$  and  $U_{z,2}$  are the vertical speeds corresponding to the two cameras). A resultant velocity magnitude was then calculated for each camera, and these two values averaged to yield a single estimate for  $U$  per experiment (Eqn 9).

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$$U = \frac{1}{2} \left( \sqrt{U_x^2 + U_y^2 + U_{z,1}^2} + \sqrt{U_x^2 + U_y^2 + U_{z,2}^2} \right)$$

Each model was sunk five times and a mean  $U$  was calculated from these replicates. Replicates beyond a threshold of  $\pm 5\%$  of the median sinking velocity were discarded from this average. Each model was dropped one additional time and photographed using a Canon 1200D DSLR camera (Tokyo, Japan) mounted on a tripod close to the tank, to obtain high resolution (18 megapixels) images which were used to determine model orientation (and thus  $L$  and  $A$ ) during settling (**Error! Reference source not found.D**).

## Wall effects

At low  $Re$ , the effects of artificial walls in an experimental (or computational) system can be nonintuitively large and lead to substantial errors if not accounted for (Vogel, 1994). Acting as an additional source of drag, the walls several tens of particle diameters away can slow down a sinking particle and increase its apparent drag coefficient. We designed our experiments to minimise wall effects by using an 0.8 m diameter tank (**Error! Reference source not found.**A) and model diameters on the order of 1 cm. To reduce potential errors further, we applied the method of Fayon & Happel (1960; summarised in Clift et al., 1978) to convert between the apparent drag coefficient when walls are present ( $C_D^{wall}$ ) and the desired drag coefficient in an unbounded domain ( $C_D^\infty$ ):

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$$C_D^\infty \approx C_D^{walls} - \frac{24}{Re} (K(\lambda) - 1)$$

where

$$K(\lambda) = \frac{1 - 0.75857 \cdot \lambda^5}{1 - 2.1050 \lambda + 2.0865 \lambda^3 - 1.7068 \lambda^5 + 0.72603 \lambda^6}$$

Here,  $\lambda = d/D$  where  $d$  is the diameter of the sinking particle and  $D$  is the tank diameter; we take  $d = L$ . While Eqn 10 is not exact, it substantially reduces the error otherwise incurred if one were to neglect wall effects entirely. Note that Eqn 10 is only valid up to about  $Re = 50$ , beyond which different corrections can be used (Clift et al., 1978).

We applied this correction by taking any experimentally determined  $C_D$  to equal  $C_D^{walls}$ , and using  $C_D^\infty$  estimated according to Eqn 10 for subsequent calculations as detailed below. In our experiments,  $\lambda$  ranged from 0.0027 – 0.0173, yielding  $K$  between 1.0057 – 1.0377. Wall effects were therefore quite small, with  $C_D^\infty / C_D^{walls}$  ranging from 0.993 – 0.994.

## Iterative approach

### First iteration

To construct an empirical cubic spline  $C_D^E(Re)$  needed to solve for  $U^0$ , at least three experimental data points (corresponding to three scale factors) are needed. These first three  $S$  were chosen by using an existing empirical  $C_D(Re)$  relationship for a sphere, valid for  $0 < Re < 10^6$  (Morrison, 2013, Fig. 8.13, page 625):

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$$C_D^M(Re) = \frac{24}{Re} + \frac{2.6 \left(\frac{Re}{5.0}\right)}{1 + \left(\frac{Re}{5.0}\right)^{1.52}} + \frac{0.411 \left(\frac{Re}{263,000}\right)^{-7.94}}{1 + \left(\frac{Re}{263,000}\right)^{-8.00}} + \left(\frac{Re^{0.80}}{461,000}\right)$$

While morphologically complex particles such as foraminifera tests (Table 1) are not expected to behave like ideal spheres, Eqn 11 should be sufficient to provide initial guesses, after which we iterate to find the solution. We note that if the particle shapes of interest were all most similar to some other

well-studied geometry (e.g., cylinders, discs, etc), using a known  $C_D(Re)$  relationship for that shape could provide better initial guesses and faster convergence.

Substituting Eqn 11 into Eqn 7 (with  $S = 1$ ,  $V = V^O$ , and  $\rho_{particle}^O, \rho_{fluid}^O, \mu^O$  substituted) and moving all terms to one side, we can numerically solve (MATLAB: *fzero* function) for our first estimate of the operating  $Re^O$ . Substituting this  $Re$  back into Eqn 7 or Eqn 11 yields an estimate of the operating  $C_D^O$ . We aimed to reproduce this  $Re$  and  $C_D$  in the first experiment, excepting that we accounted for wall effects by distinguishing between  $C_D^\infty$  and  $C_D^{walls}$  expected to occur in the tank. Hence, we could again substitute this  $Re$  into Eqn 7 but now with  $\rho_{particle}$  corresponding to the resin model and  $\rho_{fluid}$  and  $\mu$  corresponding to mineral oil, and combine this expression with Eqn 10, assuming our estimated  $C_D = C_D^\infty$ ,  $C_D^F = C_D^{walls}$ , and  $\lambda = S L^O / D$ . The resulting expression can be solved numerically for the first scale factor, termed  $S_1$ . Two more scale factors ( $S_2$  and  $S_3$ ), one smaller and one larger than  $S_1$ , were chosen to span expected  $Re$  values for forams from published literature (e.g. Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014) as well as  $Re^O$  for other species which had reached convergence. This procedure was intended to bound the correct  $S$  value that reproduces the operating  $Re^O$  and  $C_D^O$  of the settling particle. The three models were printed, their actual volumes  $V$  measured via weighing, and their settling velocities  $U$  experimentally measured as detailed in Appendix 3.

An empirical cubic spline curve  $C_D^E(Re)$  can now be fitted (D'Errico, 2009) to these three initial ( $Re, C_D$ ) data points, constrained to be monotonically decreasing and concave up within the limits of the data to match expectations for drag on objects at low to moderate  $Re$ . Three optimally spaced spline knots were used since this yielded excellent fits to the data as the number of data points increased. These details of the spline as well as its order (i.e., cubic vs linear) are somewhat arbitrary but we ensured that our results were sufficiently converged as to be insensitive to them (see Remaining iterations).

The operating point ( $Re^O, C_D^O$ ) corresponding to the particle settling in the natural environment can be visually represented as the intersection point of the  $C_D^F(Re)$  curve defined by Eqn 7 (with  $S = 1$  and  $\rho_{particle}^O, \rho_{fluid}^O, \mu^O$ ) and the empirical  $C_D^E(Re)$  relationship based on our experimental data. Algebraically, the operating point is the solution to  $C_D^F(Re) = C_D^E(Re)$ . We solved for  $Re^O$  numerically using a root finding algorithm (MATLAB's *fzero*) on the objective function  $C_D^F(Re) - C_D^E(Re) = 0$  and then obtained  $C_D^O$  by substituting  $Re^O$  into Eqn 7. Finally,  $U^O$  was easily determined from the definition of  $Re^O$  (Eqn 1 with  $U^O, L^O$ , and  $\rho_{fluid}^O$  substituted).

Since our first three empirical data points and fitted spline  $C_D^E$  corresponded to guessed model scale factors  $S$ , our initial operating point prediction ( $Re^O, C_D^O$ ) often was not located near any of these initial points or sometimes even within the bounds of these data (in which case linear extrapolation of  $C_D^E$  was used to estimate the operating point). Therefore, to ensure the accuracy of our predicted  $U^O$ , we continued iterating with additional experiments.

#### Remaining iterations

The model scale factor for the  $N^{th}$  experiment was chosen by combining Eqns 7, 8, and 10 with  $Re = Re^O$  and  $C_D^{walls} = C_D^O$ , (from the previous iteration),  $C_D^F = C_D^\infty$ , and  $V = V^{predict}$ , and

numerically solving for  $S$ . A model close to this new scale was printed and sunk, its settling velocity  $U$  recorded and  $Re$  and  $C_D$  computed, and a more accurate spline  $C_D^E$  constructed by including this new data point. The calculation of  $(Re^O, C_D^O)$  detailed in the previous section was then repeated, yielding a more accurate operating point. Overall, the aim was to tightly bound the predicted operating point with experimental data to maximize confidence in the fitted spline in this region.

The iterative process (visualized as a flowchart, **Error! Reference source not found.B**, with a specific example of convergence given in Fig. 2 C & D) was repeated until:

- 1) the predicted operating point was not extrapolated beyond our existing data,
- 2) the variation in calculated  $U^O$  between the fitting of a linear spline and cubic spline was no greater than 5%, and
- 3) the variation between the predicted  $Re^O$  and the closest experimentally measured  $Re$  was less than 15%.

In many cases, the difference between results based on four versus three data points was very small (Fig. 2 C, D), indicating rapid convergence and the possibility of streamlining the method further in the future. Through this method we calculated the sinking velocities of 30 species of planktonic foraminifera (Table 1).

## Method Validation

Our basic methodology was first validated by 3D printing a series of spherical models (10 – 20 mm in diameter) for which the theoretical  $C_D(Re)$  relationship is already well-known. In order to achieve low density (and thus low sinking velocity and low  $Re$ ), these spheres were hollow and filled with oil via two small holes (of diameter 0.8% of the sphere diameter). Our empirically generated  $C_D^E(Re)$  curve compares favourably with the theoretical  $C_D^M(Re)$  curve (Morrison, 2010) ( $R^2=0.875$ , **Error! Reference source not found.A**), with the distance between the curves approximately constant above  $Re \approx 25$ . While the error grows larger at lower  $Re$ , we expected most foraminifera species to operate at  $Re \approx 18 – 55$  based on previous work (Berger et al., 1972; Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014).

To quantify errors in our approach even more directly, we then considered hypothetical hollow spherical particles with the same material density ( $\rho_{particle}^O$ ) as foraminifera tests and a range of sizes ( $L^O = 750 – 1150 \mu m$ , similar to the species we studied) settling in seawater. This size range corresponds to  $Re = 12 – 27$ , the area where our  $C_D^E(Re)$  curve is most divergent from  $C_D^M(Re)$ . We compared predictions of the operating  $U^O$  based on our empirical  $C_D^E(Re)$  curve versus the theoretical  $C_D^M(Re)$  curve for spheres as outlined above, substituting Eqn 11 for  $C_D^E(Re)$  in the latter case. Maximum relative error in predicted  $U^O$  was 11.5% at  $Re = 16$  (corresponding to a sphere 860  $\mu m$  in diameter) while the minimum difference was 6.5% at  $Re = 27$  (corresponding to a sphere of 1150  $\mu m$  in diameter, **Error! Reference source not found.A**). This level of error is much smaller than the variation in  $U^O$  we predicted across the 30 foraminifera species we investigated (Table 1).

## 407 Time to Terminal Velocity

408 This study was concerned with predicting steady sinking speeds, but in our experiments, each model  
 409 foraminifera took a finite amount of time to accelerate from rest at the point of release to its  
 410 terminal sinking velocity. Since this transient portion of the sinking trajectory could introduce errors  
 411 into our analysis, it is important to determine whether it affected any of our recorded data.

412 During the transient acceleration phase, Eqn 2 does not hold. Instead, we can revert to the more  
 413 general form of Newton's second law:

414 12

$$415 \quad \Sigma F = F_{drag} + F_{buoyancy} - F_{weight} = -Ma = -M \frac{dU}{dt}$$

416 where  $M = V\rho_{particle}$  is particle mass, and the acceleration  $a$  can be equated to the time derivative  
 417 of instantaneous velocity  $\frac{dU}{dt}$ . A negative sign appears on the right-hand side so that we can define  
 418 the downward movement as positive for convenience. We can then substitute expressions for each  
 419 force as before:

420 13

$$421 \quad \frac{1}{2} C_D(\Psi, Re) \rho_{fluid} U^2 A + V \rho_{fluid} g - V \rho_{particle} g = -M \frac{dU}{dt}$$

422 While thus far we have not assumed anything about the particle shape, to proceed further we require  
 423 knowledge of  $C_D(\Psi, Re)$  from vanishingly small  $Re$  (when the particle is at rest) up to the terminal  
 424 velocity. Hence we will assume a spherical particle as an approximation to the model forams, so that  
 425 Morrison's empirical equation (Eqn 11) can then be substituted for  $C_D(\Psi, Re)$ :

426 14

$$427 \quad - \left\{ \frac{1}{2} \left[ \frac{24}{Re} + \frac{2.6 \left( \frac{Re}{5.0} \right)}{1 + \left( \frac{Re}{5.0} \right)^{1.52}} + \frac{0.411 \left( \frac{Re}{263,000} \right)^{-7.94}}{1 + \left( \frac{Re}{263,000} \right)^{-8.00}} + \left( \frac{Re^{0.80}}{461,000} \right) \right] \rho_{fluid} U^2 A + V \rho_{fluid} g - V \rho_{particle} g \right\} / M = \frac{dU}{dt}$$

428 Here we have isolated  $\frac{dU}{dt}$  on the right-hand side. If the definition of  $Re = \frac{L U \rho_{fluid}}{\mu}$  is inserted into 14  
 429 (not shown for brevity), one obtains an ordinary differential equation (ODE) for the unsteady velocity  
 430  $U(t)$ . The depth of the sphere  $Z(t)$  can then be obtained by solving a second much simpler ODE:

431 15

$$432 \quad U = \frac{dZ}{dt}$$

433 Both ODEs are easily solved numerically by e.g. MATLAB's *ode45* subject to the initial conditions  
 434  $U(t = 0) = 0$  and  $Z(t = 0) = 0$ .

435 It is well known that as  $Re$  approaches zero in the limit of inertia-less Stokes flow, unsteadiness can  
 436 only occur due to time-varying boundary conditions. Thus, a microorganism that stops actively  
 437 swimming will almost instantly come to a stop, and a heavy micro-particle released from rest will

almost instantly begin sinking at its terminal velocity (Purcell, 1977). As  $Re$  increases and inertia becomes increasingly important, the transient period of acceleration becomes longer. Therefore, a reasonable worst-case to examine here is the foraminifera model that sank at the highest  $Re$ .

We found *Globorotalia (Truncorotalia) truncatulinoides* to operate at  $Re = 42$  (Table 1) but here we conservatively chose the largest scale model used to generate its  $C_D^F(Re)$  spline for which  $S = 16$  and  $Re = 90$ . Inserting this model's length  $L$ , area  $A$ , and measured volume  $V$  into 14, we obtain solutions for the time varying speed and depth of a sphere approximating this model's geometry (**Error! Reference source not found.B**). The depth corresponding to where speed equals 99.9% of the terminal velocity is approximately 4.6 cm, which is much smaller than the 19 cm between where the models were released and the edge of the cameras' fields of view for data collection. Hence, the transient acceleration of each model foraminifera should have had no effect on our data or results. Most of our models should have reached terminal velocity even sooner since they sank at lower  $Re$ , e.g. within 2.2 cm for *C. dissimilis* operating at  $Re = 36$  (Table 1).

## Results & Discussion

Here we present a novel method of determining settling speed by leveraging dynamically scaled models falling under gravity rather than being towed at a controlled speed. Applying our method to foraminifera-inspired spherical particles (**Error! Reference source not found.A**), we predict settling speeds within 11.5% of theoretical expectations (**Error! Reference source not found.E**). In **Error! Reference source not found.C & D** we present an example of convergence of our method to the operating  $Re^O$ ,  $C_D^O$ , and  $U^O$  of a typical foraminifera species. There was little variation in the number of iterations required to reach convergence (mean 4, range 3-6, see Table 1), despite the morphological complexity of some species (e.g. *Globigerinoidesella fistulosa*). We suspect the higher end of this range was due to these species having forms that were particularly challenging to clean residual resin from, or the incomplete removal of air bubbles once submerged in oil.

462 Table 1. Predicted sinking speeds  $U^0$  for the 30 species of planktonic foraminifera included in this study, with each species  
463 shown in both spiral view and 90° rotation (so that the spiral view is facing to the left). <sup>1</sup> indicates species scanned high  
464 resolution. Scans of the remaining 27 species were obtained from The University of Tohoku museum's database. Operating  
465  $Re^0$  and  $C_D^0$  predicted for each species are also presented, and for comparison, the theoretical  $C_D^M$  of a sphere at the same  
466  $Re^0$ . The number of model iterations required to achieve convergence of  $U^0$  is listed (iter). Species synonyms checked at  
467 World Foraminifera Database (Hayward et al., 2018).



Image	Species	$L^0$ ( $\mu\text{m}$ )	$U^0$ ( $\text{cm s}^{-1}$ )	$Re^0$	$C_D^0$	$C_D^M(Re^0)$	Iter	Image	Species	$L^0$ ( $\mu\text{m}$ )	$U^0$ ( $\text{cm s}^{-1}$ )	$Re^0$	$C_D^0$	$C_D^M(Re^0)$	Iter
	<i>Neogloboquadrina acostaensis</i> (Blow, 1959)	307	3.6	9.3	2.5	4.2	4		<i>Sphaeroidinella dehiscent</i> (Parker & Jones, 1865)	460	5.1	32.8	1.7	1.9	4
	<i>Globigerinella adamsi</i> (Banner & Blow, 1959)	596	3.0	9.6	2.6	4.1	4		<i>Sphaeroidinella dehiscent</i> <sup>1</sup> (Parker & Jones, 1865)	269	2.5	8.3	5.0	4.5	4
	<i>Dentoglobigerina altispira</i> (Cushman & Jarvis, 1936)	282	6.1	37.3	1.6	1.8	5		<i>Catapsydrax dissimilis</i> (Cushman & Bermúdez, 1937)	380	6.2	36.4	1.7	1.8	4
	<i>Globoturborotalita apertura</i> (Cushman, 1918)	650	2.8	7.4	4.2	4.8	3		<i>Neogloboquadrina dutertrei</i> (d'Orbigny, 1839)	259	4.0	17.6	2.0	2.8	3
	<i>Globigerina bulloides</i> (d'Orbigny, 1826)	296	3.3	13.4	2.4	3.3	5		<i>Globigerinoidesella fistulosa</i> (Schubert, 1910)	392	2.6	11.2	3.0	3.7	4
	<i>Globigerinoides conglobatus</i> <sup>1</sup> (Brady, 1879)	445	1.8	5.8	5.9	5.7	4		<i>Globorotaloides hexagonus</i> (Natland, 1938)	321	2.7	8.3	3.4	4.5	4
	<i>Globorotalia (Truncorotalia) crassaformis</i> (Galloway and Wissler, 1927)	250	3.3	9.5	2.3	4.1	4		<i>Globigerinella obesa</i> (Bolli, 1957)	539	2.8	9.4	3.2	4.1	4
	<i>Neogloboquadrina humerosa</i> (Takayanagi & Saito, 1962)	360	5.0	22.7	2.0	2.4	4		<i>Globorotalia praemenardii</i> (Cushman & Stainforth, 1945)	569	3.9	14.0	2.0	3.2	4
	<i>Globoconella inflata</i> (d'Orbigny, 1839)	422	3.9	21.4	2.5	2.5	4		<i>Globoconella puncticulata</i> (d'Orbigny in Deshayes, 1832)	269	4.0	14.5	2.8	3.1	3
	<i>Fohsella lobata</i> (Bermúdez, 1949)	312	4.2	13.1	1.5	3.3	3		<i>Fohsella robusta</i> (Bolli, 1950)	312	3.5	11.2	2.3	3.7	4
	<i>Globorotalia margaritae</i> (Bolli & Bermúdez, 1965)	258	2.9	8.1	2.8	4.6	3		<i>Dentoglobigerina tripartita</i> (Koch, 1926)	303	5.7	33.6	2.0	1.9	3
	<i>Paragloborotalia mayeri</i> (Cushman & Ellisor, 1939)	298	3.1	8.1	3.2	4.5	4		<i>Paragloborotalia siakensis</i> (LeRoy, 1939)	366	3.4	10.4	2.6	3.9	4
	<i>Globoturborotalita nepenthes</i> (Todd, 1957)	353	4.2	16.4	2.4	2.9	3		<i>Globoconella sphericomiozea</i> (Walters, 1965)	627	4.3	18.8	2.3	2.7	4
	<i>Pulleniatina obliquiloculata</i> (Parker & Jones, 1865)	350	4.1	24.6	2.1	2.3	3		<i>Globorotalia (Truncorotalia) truncatulinoides</i> (d'Orbigny, 1839)	325	6.3	42.4	1.5	1.6	6
	<i>Fohsella peripheroronda</i> (Blow & Banner, 1966)	422	3.6	10.1	2.4	4.0	3		<i>Orbulina universa</i> <sup>1</sup> (d'Orbigny, 1839)	567	1.0	3.3	10.7	8.7	6
	<i>Praeorbulina curva</i> (Blow, 1956)	433	4.8	17.8	2.2	2.8	3								

Our predicted sinking speeds of foraminifera fall within aggregated existing data for 14 species (**Error! Reference source not found.**, Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014) and compare well with known speeds for other particles of comparable size and density (e.g. faecal pellets, (Table 3, Iversen & Ploug, 2010), phytoplankton (Fig. 1 in Smayda, 1971)). However, it should be noted that our predicted speeds are higher than published values for five out of the seven foraminifera species for which direct comparisons are possible (**Error! Reference source not found.**). This could be due to our ability to observe enlarged models of sinking foraminifera more accurately compared to actual specimens, and the lack of control for wall effects in previous work, which would tend to underestimate sinking speeds. There could also be considerable natural variation, which our single specimen per species (excluding *S. dehiscens*) does not capture.

Sedimentation of microscale plankton has been measured both *in situ* (e.g. Waniek, Koeve and Prien, 2000) and in the laboratory (e.g. Smayda, 1971; Miklasz and Denny, 2010). By settling dense suspensions of microorganisms, these studies provided a population sinking rate (Bienfang, 1981) which could be two to three times lower than the settling velocity of an isolated particle in the typically dilute ocean (Miklasz & Denny 2010). Other studies have, like us, used enlarged models of microscale plankton to facilitate observations. Padisak et al. (2003) used handmade models of plankton to examine drag, but there was no attempt at accurately matching  $Re$ . Holland (2010) used mechanical pencil leads as models of sinking diatom chains, keeping  $Re < 1$  in an improvement over Padisak et al. (2003). However, neither authors calculated sinking velocities for real organisms. Our dynamic scaling approach ensures that we accurately recreate the fluid flows around settling organisms – a requirement for the correct prediction of sinking speeds. We also improve on previous methodologies by effectively eliminating wall effects, basing our models on  $\mu$ CT scans, and using inexpensive cameras to observe natural sinking orientation.

By design, our dynamic scaling approach yields an interpolated  $C_D(Re)$  curve that describes the flow dynamics (and thus sinking speeds) that would occur if various fluid and/or particle parameters were varied, offering a degree of flexibility not seen in other studies. For example, phytoplankton blooms can increase both the density and viscosity of water due to exudates (Jenkinson et al., 2015), while increasing global temperatures have the opposite effect. The density and viscosity of seawater also naturally vary with latitude. Understanding how these variations affect sinking rates can offer insights into the evolutionary pressures on plankton. Our approach also allows us to isolate the effects of shape on sinking, even across species of widely varying size, density, etc, by comparing  $C_D$  of different species all hypothetically sinking at the same  $Re$ ; a manuscript focused on such biological questions relating to foraminifera is currently in preparation. Differential settling speeds of foraminifera also

502 have implications for nutrient cycling, paleoclimate reconstruction (Kucera, 2007), and the marine  
503 calcite budget (Schiebel, 2002).

504 Our method can easily be modified to study sedimenting particles operating at any  $Re$ , providing the  
505 system's  $Re$  range can be experimentally replicated. Other sinking marine particles include diatoms  
506 ( $Re \approx 10^{-2} - 1$ , Botte, Ribera D'Alcalà and Montresor, 2013) and radiolaria ( $Re \approx 10 - 200$ , Takahashi and  
507 Honjo, 1983), for which one could use digital models as we have in conjunction with a suitably viscous  
508 fluid (high viscosity silicone oil, see SI Further Applications) to enable sufficiently large models to be  
509 produced (25 cm, see SI: Further Applications). The method can also be applied to terrestrial systems  
510 such as settling spores ( $Re \approx 50$  e.g. Noblin, Yang and Dumais, 2009) and dispersing seeds ( $Re \approx 10^3$   
511 Azuma and Yasuda, 1989), again by using 3D printed models based on (often existing)  $\mu$ CT data.

512 Whilst our method pertains to settling in a quiescent fluid, one could conduct similar experiments  
513 using a flume to calculate threshold resuspension velocities (i.e. the horizontal flow speed required to  
514 lift a particle off the substrate), important in the study of wind erosion and particle transport and  
515 deposition (Bloesch, 1995; Bagnold, 1971). Similarly, studying particles suspended in shear flow could  
516 be achieved using a treadmill-like device (e.g. Durham et al (2009)) or a Taylor-Couette apparatus (e.g.  
517 Karp-Boss & Jumars (1998)). While additional dimensionless groups beyond  $Re$  and  $C_D$  would need to  
518 be matched to achieve similitude in these systems, we hope that our study provides a starting point  
519 for the experimental study of these and other more complex problems.

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## Competing interests

The authors have no competing interests to declare.

## Author contributions

SH & MW were responsible for conceptualisation of the project, MW, SH and RS contributed to developing the methodology, MW collected data and JUH & FW scanned the three foraminifera at high resolution (See Table 1). RS & MW were responsible for MATLAB code, data curation and visualisation, formal validation of the method and the initial draft of the manuscript. TH segmented the high resolution scans, generated models used for 3D printing, and assisted with 3D printing of all models. MW, SH & RS contributed to editing and reviewing the manuscript, FW & JUH provided details on methodology of high resolution scanning. SH & RS supervised MW for the duration of the project, and SH provided funding, resources and administration for the project.

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# Figure & Table Legends

## Figure 1

Fig. 1.A) Diagram of relevant forces and parameters between the model (left) and real life (right).  
B) Summary of the full method; details are discussed in main text. Boxes with thicker lines represent a decision, square boxes are data inputs, rounded square boxes are manual processes, and circles are computational steps.

## Figure 2

Fig. 2 A) Comparison of our empirically generated  $C_D^E(Re)$  curve for 3D printed spheres versus the theoretical  $C_D^M(Re)$  curve (Morrison, 2010). Goodness of fit of  $C_D^E(Re)$  to  $C_D^M(Re)$  is  $R^2 = 0.857$ . Sphere diameter is indicated for each model.  
C & D) An example of our iterative solution process for *C. dissimilis* showing best estimates of operating values (including the required model scale  $S$  to achieve similitude) based on experimental data from 3 (C), vs 4 (D) models. For reference, the theoretical  $C_D^M(Re)$  curve (Morrison, 2010) for a sphere is also shown. In C,  $S$  corresponding to the operating point is estimated as 13.84. After an additional model was sunk at  $S = 13$  (D), slightly more accurate estimates of the operating  $S$ ,  $Re^O$ , and  $C_D^O$  were obtained. When scaling the model for 3D printing only 1 decimal place was used rather than the 2 shown.  
D-I) Models of foraminifera 3D printed in clear and black resin. Models were used for public engagement but demonstrate the fidelity of the printer compared to the scan data shown Table 1.

## Table 1

Table 1. Predicted sinking speeds  $U^O$  for the 30 species of planktonic foraminifera included in this study, with each species shown in both spiral view and 90° rotation (so that the spiral view is facing to the left). <sup>1</sup> indicates species scanned high resolution. Scans of the remaining 27 species were obtained from The University of Tohoku museum's database. Operating  $Re^O$  and  $C_D^O$  predicted for each species are also presented, and for comparison, the theoretical  $C_D^M$  of a sphere at the same  $Re^O$ . The number of model iterations required to achieve convergence of  $U^O$  is listed (iter). Species synonyms checked at World Foraminifera Database (Hayward et al., 2018).